# A MESHFREE COLLOCATION METHOD FOR THE SOLUTION OF A CLASS OF PARTIAL INTEGRO-DIFFERENTIAL EQUATIONS 

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ABSTRACT: A meshfree collocation method using radial basis functions is developed for numerical solution of a parabolic type partial integro-differential equation with a weakly singular kernel. The scheme is constructed by approximating the time derivative using forward and central difference formulae while the spatial derivative is approximated using radial basis functions. Three types of radial basis functions are used for this purpose. Three test problems are provided to validate the proposed scheme. Numerical results are obtained using various numbers of collocation points and time step sizes. Accuracy of the method is assessed in terms of $L_{\infty}$ and $L_{2}$ error norms. Remarkable accuracy is obtained and the results are also compared with cubic B-spline collocation method.

Keywords: Partial integro-differential equation, Weakly singular kernel, Radial basis function, Collocation method.

## 1. INTRODUCTION

Partial integro-differential equations (PIDEs) arise as mathematical models of various physical processes including heat conduction, viscoelastic polymers, nuclear reactor dynamics, immunology, option pricing and electricity swaptions. Some applications of PIDE as models of different physical phenomena can be found in the literature [1-6]. Several numerical techniques have been developed in the literature for the approximate solution of various types of PIDEs including finite difference methods, finite element methods, wavelet methods, spline methods, chebyshev polynomials, variational iteration and homotopy perturbation methods and radial basis functions methods (see [7-23, 3132] and the references therein).

In this paper, we develop a meshfree method based on collocation principle along with radial basis functions for solution of the following type PIDEs:

$$
\int_{t>0}^{t} \beta(t-s) w_{t}(x, s) d s-w_{x x}(x, t)=g(x, t), x \in[a, b]
$$

with initial condition

$$
\begin{equation*}
w(x, 0)=h_{0}(x), \quad a \leq x \leq b, \tag{2}
\end{equation*}
$$

and boundary conditions

$$
\begin{equation*}
w(a, t)=h_{1}(t), w(b, t)=h_{2}(t), t \geq 0, \tag{3}
\end{equation*}
$$

where $\beta(t)$ is a weakly singular kernel which is given as $\beta(t)=\frac{t^{(\alpha-1)}}{\Gamma(\alpha)}, 0<\alpha<1$, at $t=0, \quad \Gamma$ denotes the gamma function, $h_{0}(x), h_{1}(t)$ and $h_{2}(t)$ are known functions.

Meshfree methods using radial basis functions (RBF) have become popular in approximation theory because of its simplicity, exponential accuracy, flexibility with respect to geometry, dimensional independence and straightforward implementation. Meshfree methods do not require a grid and only make use of a set of scattered data points regardless of the connectivity information between the points. Unlike finite-difference, finite element, finite volume and spectral methods these methods circumvent mesh generation which is main problem in mesh-based numerical methods. Also radial basis function based methods are domain type methods because the solution of the problem can be extended to whole domain In 1990, Kansa pioneered the radial basis function collocation method for the approximate solution of PDE [24]. Later on, Franke and Schaback gave a theoretical establishment to the method for the solution of partial
differential equations [5]. Further applications of this method to partial differential equations can be found in [26-30]. Recently, the radial basis functions collocation method has been used for the solution of different types of PIDEs including nonlinear parabolic partial integro-differential equation with weakly singular kernel [22], nonlinear volterra partial integro-differential equations [31] and convectiondiffusion integro-differential equations [32].

Radial basis functions are mainly divided in two categories:
(i) Infinitely smooth RBFs
(ii) Piecewise smooth RBFs

The infinitely smooth RBFs contain a parameter called shape parameter which affects both accuracy of solution and conditioning of collocation matrix, while piecewise smooth RBFs are free of shape parameter. To find that value of the shape parameter which leads to optimal accuracy is still an open problem (see [28-30]).
Some of commonly used RBFs are as follows:

| Multiquadric (MQ) | $\psi(r)=\sqrt{r^{2}+c^{2}}$ |
| :--- | :--- |
| Gaussian (GA) | $\psi(r)=\exp \left(-c r^{2}\right)$ |
| Thin plate spline (TPS) | $\psi(r)=r^{2} \ln r$ |
| Quintic | $\psi(r)=r^{5}$ |
| Spline of degree seven (SD) | $\psi(r)=r^{7}$ |

Where $c$ is the shape parameter.
In this paper, MQ, GA and SD types of radial basis functions are used for computation.

## 2. CONSTRUCTION OF PROPOSED METHOD

Consider the partial integro-differential equation given in Eqs. (1)-(3). Let $t_{n}=n k$, where $k$ is the time step and $t_{n+1}=t_{n}+k, n=0,1,2 \ldots$. In Eq. (1), the integral term is approximated at $t=t_{n+1}$ as given in [19, 23]:
$\int_{0}^{t_{n+1}} \frac{\left(t_{n+1}-s\right)^{\alpha-1} w_{t}(x, s) d s}{\Gamma(\alpha)}=\int_{t_{0}}^{t_{1}} \frac{\left(w^{1}(x)-w^{0}(x)\right)\left(t_{n+1}-s\right)^{\alpha-1} d s}{k \Gamma(\alpha)}+$ $\sum_{r=1}^{n} \int_{t_{r}}^{t_{r+1}} \frac{\left(w^{r+1}(x)-w^{r-1}(x)\right)\left(t_{n+1}-s\right)^{\alpha-1} d s}{2 k \Gamma(\alpha+1)}=b_{n} \frac{w^{1}(x)-w^{0}(x)}{2 \Gamma(\alpha+1) k^{1-\alpha}}+$
$\frac{1}{2 \Gamma(\alpha+1)} \sum_{r=0}^{n-1} b_{r} \frac{w^{r+1}(x)-w^{r-1}(x)}{k^{1-\alpha}}$,
where $w^{n+1}(x)=w\left(x, t_{n+1}\right), b_{n}=(n+1)^{\alpha}-n^{\alpha}, n=$ $0,1,2, . ., M$.
From Eqs. (1) and (4), we have
$b_{0} w^{n+1}(x)-2 \Gamma(\alpha+1) k^{1-\alpha} w_{x x}^{n+1}(x)=b_{0} w^{n-1}(x)-$
$\sum_{r=1}^{n-1} b_{r}\left(w^{n-r+1}(x)-w^{n-r-1}(x)\right)-$
$2 b_{n}\left(w^{1}(x)-w^{0}(x)\right)+2 \Gamma(\alpha+1) k^{1-\alpha} g^{n+1}(x)$,
(5)
where $w_{x x}^{n+1}(x)=w_{x x}\left(x, t_{n+1}\right), g^{n+1}(x)=g\left(x, t_{n+1}\right)$,
Assuming $d_{0}=2 \Gamma(\alpha+1) k^{1-\alpha}$ with $b_{0}=1$.
Eq. (5) becomes
$w(x)-d_{0} w_{x x}^{n+1}(x)=$
$-b_{1} w^{n}(x)+\sum_{r=1}^{n-1}\left(b_{r-1}-b_{r+1}\right) w^{n-r}(x)-b_{n} w^{1}(x)+$
$\left(b_{n-1}+2 b_{n}\right) w^{0}(x)+d_{0} g^{n+1}(x), \quad n \geq 1$.
(6)

Substituting $n=0$, in Eq. (6), then $u^{1}$ can be obtained from the following:

$$
w^{1}(x)-0.5 d_{0} w_{x x}^{1}=w^{0}(x)+0.5 d_{0} g^{1}(x)
$$

To obtain approximate solution of the problem (1)-(3) using radial basis functions collocation method, we choose $N+1$ distinct points $x_{i}, i=0,1,2, \ldots, N$, from the interval [ $a, b]$ such that $a=x_{0}$ and $x_{N}=b$. The RBF approximation of the function $w(x, t)$ at $n$th time level is given by
$w^{n}(x)=\sum_{m=0}^{N} \lambda_{m}^{n} \psi\left(e_{m}\right)$,
where $\lambda_{m}^{n}$ are unknown time dependent parameters to be determined from collocation conditions, $\psi$ is a radial basis function and $e_{m}=\left|x-x_{m}\right|$ is Euclidean distance. Collocating Eq. (6) at $x=x_{i}, i=1,2, \ldots, N-1$, and using Eq. (7), we get
$\sum_{m=0}^{N} \lambda_{m}^{n+1} \psi\left(e_{i m}\right)-$
$d_{0} \sum_{m=0}^{N} \lambda_{m}^{n+1} \psi^{\prime \prime}\left(e_{i m}\right)=$
$-b_{1} \sum_{m=0}^{N} \lambda_{m}^{n} \psi\left(e_{i m}\right)+\sum_{r=1}^{n-1}\left(\left(b_{r-1}-\right.\right.$
$\left.\left.b_{r+1}\right) \sum_{m=0}^{N} \lambda_{m}^{n-r} \psi\left(e_{i m}\right)\right)-b_{n} \sum_{m=0}^{N} \lambda_{m}^{1} \psi\left(e_{i m}\right)+$
$\left(b_{n-1}+2 b_{n}\right) \sum_{m=0}^{N} \lambda_{m}^{0} \psi\left(e_{i m}\right)+d_{0} g^{n+1}\left(x_{i}\right)$.
Also from Eqs. (3) and (7), we have
$\left.\sum_{m=0}^{N} \lambda_{m}^{n+1} \psi\left(e_{0 m}\right)=h_{1}\left(t_{n+1}\right),\right\}$
$\sum_{m=0}^{N} \lambda_{m}^{n+1} \psi\left(e_{N m}\right)=h_{2}\left(t_{n+1}\right)$,
where $e_{i m}=\left|x_{i}-x_{m}\right|$.
In matrix form Eqs. (8) and (9) can be expressed as $\lambda^{n+1}=\mathbf{C}^{-1}\left[-b_{1} \mathbf{A} \lambda^{n}+\sum_{r=1}^{n-1}\left(b_{r-1}-b_{r+1}\right) \mathbf{A} \lambda^{n-\mathrm{r}}-b_{n} \mathbf{A} \boldsymbol{\lambda}^{1}+\right.$ $\left.d_{0} \boldsymbol{G}^{n+1}\right]$,
(10)
where $\mathbf{A}=\left[a_{i j}\right]_{i, j=0}^{N}, \mathbf{B}=\left[b_{i j}\right]_{i, j=0}^{N}$ such that $a_{i j}=\psi\left(e_{i j}\right)$, $b_{i j}=\left\{\begin{array}{c}\psi^{\prime \prime}\left(e_{i j}\right), 1 \leq i \leq N-1,0 \leq j \leq N, \\ 0, \quad i=0, N, 0 \leq j \leq N .\end{array}\right.$
$\mathbf{C}=\left[\mathbf{A}-d_{0} \mathbf{B}\right]$ and $\lambda^{n}=\left[\lambda_{0}^{n}, \lambda_{1}^{n}, \ldots, \lambda_{N}^{n}\right]^{T}$. Inevitability of the matrix $\mathbf{C}$ is yet to be proved. Eq. (10) represents a system of $N+1$ equations in $N+1$ unknown parameters $\lambda_{j}^{n+1}$ 's. The solution of Eq. (10) leads to the vector $\lambda^{n+1}$ and then the approximate solution is obtained by using Eq. (7).

## 3. NUMERICAL TESTS AND PROBLEMS

In this section we provide some problems in order to test accuracy of the RBF collocation method (10) for the solution of the problem (1)-(3). For the sake of comparison all the three test problems are taken from the reference [19] with $\alpha=\frac{1}{2} . \quad$ Error norms $L_{\infty}$ and $L_{2}$ are used for this purpose. Accuracy of the proposed scheme is compared with results of cubic B-spline collocation method [19]. The solution plotted in each figure is obtained using multiquadric in the RBF collocation method (10).

Example 1: Consider Eqs. (1)-(3) with $x \in[0,1]$, and $g(x, t)$ is chosen so that the exact solution [19] is $w(x, t)=(t+1) \sin \pi x$.
The initial and boundary conditions are obtained from the exact solution. Computations are performed using MQ, SD and GA types of RBFs with parameters $N=60,100, k=$ $0.0001,0.001$ and the error norms $L_{\infty}$ and $L_{2}$ up to time level $M=1000$ are reported in Tables (1)-(2) whereas Table (3) contains the error norms for different number of collocation points $N=10,20,30,40,60$. Better accuracy of the present method using MQ and GA than cubic B-spline collocation method [19] is obvious from Tables 1-2. However, SD produces better results than [19] for larger values of $N$. Fig. 1 presents the RBF and exact solutions up to $M=1000$. Fig. 2 shows the RBF solutions over the time interval [0, 1]. Fig. 3 shows $L_{2}$ errors versus $1 / N$. Most accurate results are obtained for shape parameter $c=$ $0.21,0.125$ corresponding to $N=60,100$ respectively for MQ and $c=100$ for GA while SD is free of shape parameter $c$.
Example 2: In this example we consider Eqs. (1)-(3) and choose $g(x, t)$ such that the exact solution [19] is
$w(x, t)=(t+1) \cos \pi x, x \in[0,1]$.
The initial and boundary conditions are taken from the exact solution. Numerical simulations are done using MQ, SD and GA with parameters $N=60, k=0.001,0.0001$ and the error norms $L_{\infty}$ and $L_{2}$ up to time level $M=1000$ are reported in Table (4) while Table 5 contains the error norms for different number of collocation points $N=10,20,30$, 40, 50, 60. Better accuracy of the present method using MQ, SD and GA than B-spline collocation methods [19, 23] is evident from Table (4). However, SD requires larger values of $N$ for better accuracy than [19, 23]. Fig. 4 presents the RBF and exact solutions up to time level $M=500$. Fig. 5 shows the RBF solutions over the time interval [0, 1]. Most accurate results are obtained for shape parameter $c=0.21$ for MQ and $c=80,90$ corresponding to $k=0.001,0.0001$ respectively for GA. The radial basis functions SD is independent of the shape parameter $c$.
Example 3: We consider Eqs. (1)-(3) and choose $g(x, t)$ such that the exact solution [19] is
$w(x, t)=(t+1)^{2} \sin \pi x, x \in[-1,1]$.
The initial and boundary conditions are taken from the exact solution. Simulations are done using MQ, SD and GA with parameters $N=40, k=0.001,0.00125$ and the error norms $L_{\infty}$ and $L_{2}$ up to time level $M=500$ are reported in Table (6) Better accuracy of the present method using MQ, SD and GA than cubic B-spline collocation method [19] can be seen from Table (6). However, SD requires larger values of $N$ for better accuracy than [19]. Fig. 6 presents the RBF and exact solutions up to time level $M=1000$. Fig. 7 shows the RBF solutions over the time interval [0,1]. Most accurate results are obtained for shape parameter $c=0.4,0.32$ corresponding to $k=0.001,0.00125$ respectively for MQ and $\quad c=75,70 \quad$ corresponding to $k=0.001,0.00125$ respectively for GA. The radial basis function SD is independent of shape parameter $c$.

Table 1: $L_{\infty}$ and $L_{2}$ for $\boldsymbol{k}=0.0001$.

| $N=60$ |  |  |  | $N=100$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | M | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |
| $\begin{aligned} & \mathrm{M} \\ & \mathrm{Q} \end{aligned}$ | 10 | $1.67 \times 10^{-7}$ | $8.48 \times 10^{-9}$ | $1.66 \times 10^{-7}$ | $7.13 \times 10^{-9}$ |
|  | 50 | $1.74 \times 10^{-7}$ | $1.20 \times 10^{-8}$ | $1.71 \times 10^{-7}$ | $9.65 \times 10^{-9}$ |
|  | 100 | $1.79 \times 10^{-7}$ | $1.43 \times 10^{-8}$ | $1.73 \times 10^{-7}$ | $1.07 \times 10^{-8}$ |
|  | 500 | $1.87 \times 10^{-7}$ | $1.89 \times 10^{-8}$ | $1.82 \times 10^{-7}$ | $1.38 \times 10^{-8}$ |
|  | 1000 | $1.98 \times 10^{-7}$ | $2.07 \times 10^{-8}$ | $1.91 \times 10^{-7}$ | $1.55 \times 10^{-8}$ |
| SD | 10 | $1.01 \times 10^{-5}$ | $5.54 \times 10^{-7}$ | $1.22 \times 10^{-6}$ | $5.11 \times 10^{-8}$ |
|  | 50 | $1.04 \times 10^{-5}$ | $7.51 \times 10^{-7}$ | $1.25 \times 10^{-6}$ | $6.88 \times 10^{-8}$ |
|  | 100 | $1.06 \times 10^{-5}$ | $8.52 \times 10^{-7}$ | $1.26 \times 10^{-6}$ | $7.79 \times 10^{-8}$ |
|  | 500 | $1.12 \times 10^{-5}$ | $1.10 \times 10^{-6}$ | $1.33 \times 10^{-6}$ | $1.00 \times 10^{-7}$ |
|  | 1000 | $1.22 \times 10^{-5}$ | $1.22 \times 10^{-6}$ | $1.40 \times 10^{-6}$ | $1.12 \times 10^{-7}$ |
| GA | 10 | $3.85 \times 10^{-7}$ | $2.21 \times 10^{-8}$ | $\begin{aligned} & 8.45 \times \\ & 10^{-7} \end{aligned}$ | $2.98 \times 10^{-8}$ |
|  | 50 | $3.94 \times 10^{-7}$ | $2.98 \times 10^{-8}$ | $\begin{aligned} & 7.91 \times \\ & 10^{-7} \end{aligned}$ | $3.50 \times 10^{-8}$ |
|  | 100 | $4.23 \times 10^{-7}$ | $3.46 \times 10^{-8}$ | $\begin{aligned} & 9.13 \times \\ & 10^{-7} \end{aligned}$ | $5.09 \times 10^{-8}$ |
|  | 500 | $4.56 \times 10^{-7}$ | $4.53 \times 10^{-8}$ | $\begin{aligned} & 9.51 \times \\ & 10^{-7} \end{aligned}$ | $6.38 \times 10^{-8}$ |
|  | 1000 | $4.82 \times 10^{-7}$ | $4.98 \times 10^{-8}$ | $\begin{aligned} & 9.82 \times \\ & 10^{-7} \end{aligned}$ | $7.20 \times 10^{-8}$ |
| $\begin{aligned} & {[19} \\ & ] \end{aligned}$ | 10 | $4.12 \times 10^{-5}$ | $3.76 \times 10^{-6}$ |  |  |
|  | 50 | $4.06 \times 10^{-6}$ | $3.71 \times 10^{-7}$ |  |  |
| Table 2: $L_{\infty}$ and $L_{2}$ for $k=0.001$. |  |  |  |  |  |
| $N=60$ |  |  |  | $N=100$ |  |
|  | M | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |
| $\begin{aligned} & \mathrm{M} \\ & \mathrm{Q} \end{aligned}$ | 10 | $2.88 \times 10^{-8}$ | $2.66 \times 10^{-9}$ | $1.13 \times 10^{-7}$ | $6.85 \times 10^{-9}$ |
|  | 50 | $4.88 \times 10^{-8}$ | $4.63 \times 10^{-9}$ | $1.17 \times 10^{-7}$ | $9.03 \times 10^{-9}$ |
|  | 100 | $4.71 \times 10^{-8}$ | $4.70 \times 10^{-9}$ | $1.29 \times 10^{-7}$ | $1.04 \times 10^{-8}$ |
|  | 500 | $6.67 \times 10^{-8}$ | $7.29 \times 10^{-9}$ | $1.76 \times 10^{-7}$ | $1.53 \times 10^{-8}$ |
|  | 1000 | $7.47 \times 10^{-8}$ | $7.74 \times 10^{-9}$ | $2.40 \times 10^{-7}$ | $2.17 \times 10^{-8}$ |
| SD | 10 | $1.06 \times 10^{-5}$ | $8.13 \times 10^{-7}$ | $1.26 \times 10^{-6}$ | $7.44 \times 10^{-8}$ |
|  | 50 | $1.23 \times 10^{-5}$ | $1.08 \times 10^{-6}$ | $1.33 \times 10^{-6}$ | $9.87 \times 10^{-8}$ |
|  | 100 | $1.18 \times 10^{-5}$ | $1.21 \times 10^{-6}$ | $1.40 \times 10^{-6}$ | $1.10 \times 10^{-7}$ |
|  | 500 | $1.62 \times 10^{-5}$ | $1.80 \times 10^{-6}$ | $1.91 \times 10^{-6}$ | $1.64 \times 10^{-7}$ |
|  | 1000 | $2.17 \times 10^{-5}$ | $2.45 \times 10^{-6}$ | $2.56 \times 10^{-6}$ | $2.23 \times 10^{-7}$ |
| GA | 10 | $4.63 \times 10^{-7}$ | $2.26 \times 10^{-8}$ | $5.30 \times 10^{-8}$ | $2.72 \times 10^{-9}$ |
|  | 50 | $3.74 \times 10^{-7}$ | $2.75 \times 10^{-8}$ | $1.80 \times 10^{-7}$ | $8.20 \times 10^{-9}$ |
|  | 100 | $3.06 \times 10^{-7}$ | $2.52 \times 10^{-8}$ | $1.20 \times 10^{-7}$ | $5.89 \times 10^{-9}$ |
|  | 500 | $4.63 \times 10^{-7}$ | $4.50 \times 10^{-8}$ | $2.95 \times 10^{-7}$ | $1.78 \times 10^{-8}$ |
|  | 1000 | $5.86 \times 10^{-7}$ | $6.08 \times 10^{-8}$ | $3.31 \times 10^{-7}$ | $2.29 \times 10^{-8}$ |
| $\begin{aligned} & {[19} \\ & ] \end{aligned}$ | 10 | $8.77 \times 10^{-4}$ | $8.01 \times 10^{-5}$ |  |  |
|  | 50 | $8.23 \times 10^{-4}$ | $7.51 \times 10^{-5}$ |  |  |



Figure 1: RBF and exact solutions for Example-1 for $\quad k=$ 0.001.


Figure 2: RBF solutions over time interval $[0,1]$ corresponding to Example-1 for $N=60, k=0.001$.

Table 3: $L_{\infty}$ and $L_{2}$ for $k=0.0001, M=10$.

| MQ |  |  | [19] |  |
| :---: | :---: | :---: | :---: | :---: |
| $N$ | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |
| 10 | $9.33 \times 10^{-8}$ | $\begin{aligned} & 1.44 \\ & \times 10^{-8} \end{aligned}$ | $8.58 \times 10^{-4}$ | $\begin{aligned} & 1.84 \times \\ & 10^{-4} \end{aligned}$ |
| 20 | $3.18 \times 10^{-7}$ | $3.61 \times 10^{-8}$ | $1.62 \times 10^{-4}$ | $\begin{aligned} & 2.06 x \\ & 10^{-5} \end{aligned}$ |
| 30 | $1.14 \times 10^{-7}$ | $9.40 \times 10^{-9}$ | $4.09 \times 10^{-5}$ | $\begin{aligned} & 2.96 x \\ & 10^{-6} \end{aligned}$ |
| 40 | $4.14 \times 10^{-9}$ | $\begin{aligned} & 2.88 \times \\ & 10^{-10} \end{aligned}$ | $1.32 \times 10^{-6}$ | $\begin{aligned} & 1.16 x \\ & 10^{-6} \end{aligned}$ |
| 60 | $1.67 \times 10^{-7}$ | $8.48 \times 10^{-9}$ | $9.14 \times 10^{-5}$ | $\begin{aligned} & 6.85 x \\ & 10^{-6} \end{aligned}$ |
|  | GA |  | SD |  |
| 10 | $1.30 \times 10^{-8}$ | $2.77 \times 10^{-9}$ | $4.90 \times 10^{-3}$ | $\begin{aligned} & 8.01 \times \\ & 10^{-4} \end{aligned}$ |
| 20 | $6.39 \times 10^{-8}$ | $6.00 \times 10^{-9}$ | $2.14 \times 10^{-4}$ | $\begin{aligned} & 2.18 \times \\ & 10^{-5} \end{aligned}$ |
| 30 | $7.09 \times 10^{-8}$ | $7.02 \times 10^{-9}$ | $3.88 \times 10^{-5}$ | $\begin{aligned} & 3.08 \times \\ & 10^{-6} \end{aligned}$ |
| 40 | $1.55 \times 10^{-7}$ | $8.22 \times 10^{-9}$ | $1.18 \times 10^{-5}$ | $\begin{aligned} & 7.93 \times \\ & 10^{-7} \end{aligned}$ |
| 60 | $3.85 \times 10^{-7}$ | $2.21 \times 10^{-8}$ | $2.34 \times 10^{-6}$ | $\begin{aligned} & 1.48 \times \\ & 10^{-7} \\ & \hline \end{aligned}$ |


| Table 4: $\boldsymbol{L}_{\infty}$ and $\boldsymbol{L}_{\mathbf{2}}$ for $\boldsymbol{N}=\mathbf{6 0}$ |  |  |  |  |  |  |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  | SD | $[19]$ |  |
| $N$ | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |  |  |  |  |
| 10 | $4.29 \times 10^{-2}$ | $7.20 \times 10^{-3}$ | $2.02 \times 10^{-3}$ | $4.51 \times 10^{-4}$ |  |  |  |  |
| 20 | $1.20 \times 10^{-3}$ | $1.25 \times 10^{-4}$ | $4.29 \times 10^{-4}$ | $6.78 \times 10^{-5}$ |  |  |  |  |
| 30 | $1.92 \times 10^{-4}$ | $1.57 \times 10^{-5}$ | $1.35 \times 10^{-4}$ | $1.74 \times 10^{-5}$ |  |  |  |  |
| 40 | $5.55 \times 10^{-5}$ | $3.84 \times 10^{-6}$ | $3.22 \times 10^{-5}$ | $3.60 \times 10^{-6}$ |  |  |  |  |
| 50 | $2.16 \times 10^{-5}$ | $1.32 \times 10^{-6}$ | $1.54 \times 10^{-5}$ | $1.54 \times 10^{-6}$ |  |  |  |  |
| 60 | $1.01 \times 10^{-5}$ | $5.54 \times 10^{-7}$ | $4.12 \times 10^{-5}$ | $3.76 \times 10^{-6}$ |  |  |  |  |
| $\times 10^{-4}$ |  |  |  |  |  |  |  |  |



Figure 3: $L_{2}$ error versus 1/N Example-1 for $k=0.001$.


Figure 4: RBF and exact solutions corresponding to le-2 for $N=60, k=0.001$.

Table 5: $L_{\infty}$ and $L_{2}$ for $k=0.0001, M=10$.

| Table 5: $\boldsymbol{L}_{\infty}$ and $\boldsymbol{L}_{\mathbf{2}}$ for $\boldsymbol{k}=\mathbf{0} \mathbf{0 0 0 1}, \boldsymbol{M}=\mathbf{1 0}$ |  |  |  |  |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :---: |
| $k=0.001$ |  |  |  |  | $k=0.0001$ |  |
|  | $M$ | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |  |
|  | 10 | $3.06 \times 10^{-7}$ | $2.00 \times 10^{-8}$ | $2.43 \times 10^{-7}$ | $9.18 \times 10^{-9}$ |  |
| M | 50 | $3.18 \times 10^{-7}$ | $2.12 \times 10^{-8}$ | $4.14 \times 10^{-7}$ | $2.03 \times 10^{-8}$ |  |
| Q | $3.25 \times 10^{-7}$ | $2.25 \times 10^{-8}$ | $4.10 \times 10^{-7}$ | $1.64 \times 10^{-8}$ |  |  |
|  | 100 | $3.39 \times 10^{-7}$ | $2.38 \times 10^{-8}$ | $3.70 \times 10^{-7}$ | $1.86 \times 10^{-8}$ |  |
|  | 500 | $4.62 \times 10^{-7}$ | $3.32 \times 10^{-8}$ | $3.57 \times 10^{-7}$ | $2.12 \times 10^{-8}$ |  |
|  | 10 | $2.34 \times 10^{-6}$ | $1.48 \times 10^{-7}$ | $2.25 \times 10^{-6}$ | $1.20 \times 10^{-7}$ |  |
|  | 20 | $2.38 \times 10^{-6}$ | $1.56 \times 10^{-7}$ | $2.28 \times 10^{-6}$ | $1.32 \times 10^{-7}$ |  |
| SD | 50 | $2.46 \times 10^{-6}$ | $1.66 \times 10^{-7}$ | $2.32 \times 10^{-6}$ | $1.44 \times 10^{-7}$ |  |
|  | 100 | $2.59 \times 10^{-6}$ | $1.77 \times 10^{-7}$ | $2.35 \times 10^{-6}$ | $1.51 \times 10^{-7}$ |  |
|  | 500 | $3.54 \times 10^{-6}$ | $2.46 \times 10^{-7}$ | $2.47 \times 10^{-6}$ | $1.67 \times 10^{-7}$ |  |
|  | 10 | $1.55 \times 10^{-7}$ | $8.18 \times 10^{-9}$ | $2.43 \times 10^{-7}$ | $1.16 \times 10^{-8}$ |  |
|  | 20 | $1.59 \times 10^{-7}$ | $9.49 \times 10^{-9}$ | $3.22 \times 10^{-7}$ | $1.49 \times 10^{-8}$ |  |
| GA | 50 | $1.64 \times 10^{-7}$ | $8.73 \times 10^{-9}$ | $2.00 \times 10^{-7}$ | $1.16 \times 10^{-8}$ |  |


|  | 100 | $1.71 \times 10^{-7}$ | $1.07 \times 10^{-8}$ | $3.24 \times 10^{-7}$ | $1.64 \times 10^{-8}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | 500 | $2.35 \times 10^{-7}$ | $1.26 \times 10^{-8}$ | $2.33 \times 10^{-7}$ | $1.43 \times 10^{-8}$ |
| $[19$ | 10 | $9.90 \times 10^{-4}$ | $8.59 \times 10^{-5}$ | $9.14 \times 10^{-5}$ | $6.85 \times 10^{-6}$ |
|  | 20 | $9.88 \times 10^{-4}$ | $8.58 \times 10^{-5}$ | $9.10 \times 10^{-5}$ | $5.40 \times 10^{-6}$ |
|  | 50 | $9.77 \times 10^{-4}$ | $8.48 \times 10^{-5}$ | $8.86 \times 10^{-5}$ | $5.44 \times 10^{-6}$ |
| $[23$ | 10 | $7.29 \times 10^{-6}$ | $8.15 \times 10^{-7}$ | $1.02 \times 10^{-5}$ | $1.14 \times 10^{-6}$ |
|  | 20 | $1.16 \times 10^{-5}$ | $1.29 \times 10^{-6}$ | $1.02 \times 10^{-5}$ | $1.14 \times 10^{-6}$ |
|  | 50 | $2.06 \times 10^{-5}$ | $2.30 \times 10^{-6}$ | $2.90 \times 10^{-5}$ | $3.24 \times 10^{-6}$ |



Figure 5: RBF solutions over time interval [0, 1] for Example-2 for $N=60, k=0.001$.


Figure 6: RBF and exact solutions corresponding to Example-3 for $N=40, k=0.001$.

Table 6: $L_{\infty}$ and $L_{2}$ for $N=60$.

|  | $k=0.001$ |  |  | $k=0.00125$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | M | $L_{\infty}$ | $L_{2}$ | $L_{\infty}$ | $L_{2}$ |
| $\begin{aligned} & \mathrm{M} \\ & \mathrm{Q} \end{aligned}$ | 10 | $5.98 \times 10^{-6}$ | $6.69 \times 10^{-7}$ | $8.82 \times 10^{-6}$ | $9.84 \times 10^{-7}$ |
|  | 20 | $9.99 \times 10^{-6}$ | $1.12 \times 10^{-6}$ | $1.46 \times 10^{-5}$ | $1.63 \times 10^{-6}$ |
|  | 50 | $1.86 \times 10^{-5}$ | $2.08 \times 10^{-6}$ | $2.68 \times 10^{-5}$ | $3.00 \times 10^{-6}$ |
|  | 10 | $2.87 \times 10^{-5}$ | $3.21 \times 10^{-6}$ | $4.11 \times 10^{-5}$ | $4.59 \times 10^{-6}$ |
|  | 0 |  |  |  |  |
|  | 50 | $7.27 \times 10^{-5}$ | $8.13 \times 10^{-6}$ | $1.03 \times 10^{-4}$ | $1.15 \times 10^{-5}$ |
|  | 0 |  |  |  |  |
| SD | 10 | $6.06 \times 10^{-5}$ | $6.26 \times 10^{-6}$ | $6.14 \times 10^{-5}$ | $6.78 \times 10^{-6}$ |
|  | 20 | $6.31 \times 10^{-5}$ | $7.56 \times 10^{-6}$ | $6.43 \times 10^{-5}$ | $8.30 \times 10^{-6}$ |
|  | 50 | $6.86 \times 10^{-5}$ | $9.97 \times 10^{-6}$ | $7.32 \times 10^{-5}$ | $1.10 \times 10^{-5}$ |
|  | 10 | $8.22 \times 10^{-5}$ | $1.22 \times 10^{-6}$ | $9.84 \times 10^{-5}$ | $1.40 \times 10^{-5}$ |
|  | 0 |  |  |  |  |
|  | 50 | $1.86 \times 10^{-4}$ | $2.62 \times 10^{-5}$ | $2.37 \times 10^{-4}$ | $3.28 \times 10^{-5}$ |
|  | 0 |  |  |  |  |
| GA | 10 | $9.08 \times 10^{-6}$ | $5.90 \times 10^{-7}$ | $6.31 \times 10^{-6}$ | $6.52 \times 10^{-7}$ |
|  | 20 | $9.14 \times 10^{-6}$ | $6.63 \times 10^{-7}$ | $1.13 \times 10^{-5}$ | $1.16 \times 10^{-6}$ |
|  | 50 | $1.16 \times 10^{-5}$ | $1.19 \times 10^{-6}$ | $2.25 \times 10^{-5}$ | $2.36 \times 10^{-6}$ |


|  | 10 | $2.01 \times 10^{-5}$ | $2.05 \times 10^{-6}$ | $3.57 \times 10^{-5}$ | $3.81 \times 10^{-6}$ |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 0 |  |  |  |  |  |
|  | 50 | $5.46 \times 10^{-5}$ | $5.64 \times 10^{-6}$ | $9.02 \times 10^{-5}$ | $9.69 \times 10^{-6}$ |
| 0 |  |  |  |  |  |
| $[19$ | 10 | $5.99 \times 10^{-4}$ | $6.25 \times 10^{-5}$ | $1.00 \times 10^{-3}$ | $1.05 \times 10^{-4}$ |
| $]$ |  |  |  |  |  |
|  | 20 | $4.33 \times 10^{-4}$ | $4.66 \times 10^{-5}$ | $7.23 \times 10^{-4}$ | $7.71 \times 10^{-5}$ |
| 50 | $2.00 \times 10^{-3}$ | $1.67 \times 10^{-4}$ | $2.87 \times 10^{-3}$ | $2.34 \times 10^{-4}$ |  |



Figure 7: RBF solutions over time interval [0,1] for Example-3 for $N=40, k=0.001$.

## 4. CONCLUSION

A collocation method coupled with radial basis functions is employed to approximate solution of a parabolic type integro-differential equation with a weakly singular kernel. The proposed method is validated by implementing three benchmark problems from literature. The errors are satisfactorily small and the results are in good agreement with exact solution. Implementation of the method is simple like finite difference method. Infinitely RBFs provided excellent accuracy. Numerical simulations suggest that this method can be used for numerical approximation of integral equations, partial differential equations and partial integrodifferential equations of such type.

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